



What can biologists say about galaxy evolution ?

Didier Fraix-Burnet, Philippe Choler, Emmanuel J.P. Douzery

► To cite this version:

Didier Fraix-Burnet, Philippe Choler, Emmanuel J.P. Douzery. What can biologists say about galaxy evolution ?. Astrophysics and Space Science, 2003, 284, pp.535-538. 10.1007/978-94-017-3315-1_48 . hal-00000241

HAL Id: hal-00000241

<https://hal.science/hal-00000241>

Submitted on 18 Mar 2003

HAL is a multi-disciplinary open access archive for the deposit and dissemination of scientific research documents, whether they are published or not. The documents may come from teaching and research institutions in France or abroad, or from public or private research centers.

L'archive ouverte pluridisciplinaire **HAL**, est destinée au dépôt et à la diffusion de documents scientifiques de niveau recherche, publiés ou non, émanant des établissements d'enseignement et de recherche français ou étrangers, des laboratoires publics ou privés.

What can biologists say about galaxy evolution ?

Didier Fraix-Burnet

Laboratoire d'Astrophysique de Grenoble, BP 53, F-38041 Grenoble cedex 9, France

Philippe Choler

Laboratoire de Biologie des Populations d'Altitude, BP 53, F-38041 Grenoble cedex 9, France

Emmanuel Douzery

Institut des Sciences de l'Évolution de Montpellier, F-34095 Montpellier cedex 5, France

September 15, 2002

Abstract. It is possible to borrow from a topic of biology called phylogenetic systematics, concepts and tools for a logical and objective classification of galaxies. It is based on observable properties of objects - characters - either qualitative (like morphology) or quantitative (like luminosity, mass or spectrum). Distance analysis can readily be performed using a method called phenetics and based on characters. But the most promising approach is cladistics. It makes use of characters that can exist in at least two states, one being ancestral and the other one derived. Objects are gathered depending on the derived states they share. We illustrate a first application of this method to astrophysics, that we name astrocladistics, with dwarf galaxies from the Local Group.

Keywords: galaxies: classification – galaxies: evolution – cladistics

1. Introduction

Biologists have a long experience in the classification of complex objects that are in evolution. It dates back to Aristote (around 100 A.D.) who attempted a first classification of living organisms. A lot of different systems have been proposed, but a first revolution came with Linné (around 1730) and his binomial nomenclature that is still in use today! Its success is to be found in the fact that the names are unrelated to properties of objects. A few years later, Adanson (in 1763) proposed that an objective classification scheme can only be obtained by using all characters of an object. This idea has been incredibly successful and gave birth to what would be called nowadays phenetics or multivariate distance analysis. It was not until Darwin (around 1850) that evolution was found to be at the origin of the hierarchy and diversity of living organisms. It is then another century before Hennig (in 1950) devised a methodology for introducing the evolution in the classification process itself. This is called cladistics and has revolutionized somewhat the



evolutionary classification of species (see an interesting discussion in Stewart 1993).

Systematics (science of classification of living organisms) distinguishes three categories of classification: 1) from apparent look (based on a few characters), 2) from global similarity (based on all characters), 3) from common history (based on the evolution of characters). Until the XVIIIth century, biologists classified living organisms according to the first category, the traditional way. It consists of choosing one or two characters (fruits, flowers, legs, wings) depending on very subjective a priori assumptions. It yielded quite a messy picture of Nature, with organisms not fitting in a given classification system and even incompatibilities between different systems. The second category is phenetics and benefitted a lot from the Linné's binomial nomenclature. It is not clear who first proposed a hierarchical organization for living species, but Linné clearly defined the two names as to correspond respectively to *genus* and *species*. This produced the well known evolutionary trees. Nowadays, these trees are mainly a result of cladistics analysis, that corresponds to the last category of classification. It does not compare the values of characters, like in phenetics, but compare the states of evolution of each character. This information can be gained from observations (paleontology), from models or theories, or from comparison with other groups of objects. In a sense, cladograms (trees obtained from cladistics) are visualizations of the evolutionary schemes hypothesized for all characters.

Since Hubble in the 1930s, galaxies have always been classified according to the traditional way (e.g. Roberts & Haynes 1994 and references therein). Nomenclature heavily follows physical properties (morphology, size, activity, luminosity at a given wavelength, radio loudness, etc...). But an overwhelming amount of new data compels us to consider galaxies as very complex and diverse objects, and their history as a quite complicated puzzle. We propose that a lot can be gained by paralleling the extragalactic context with biology. Some fundamental and fascinating questions can be raised: how do we describe a galaxy? what is the life of a galaxy? what is the history of galaxies considered as a species? what about the environment? Following Linné's approach, it seems to us that a new nomenclature will have to be coined for galaxies, based on a much more objective classification.

2. Astrocladistics

Multivariate distance analysis (second category of classification) can easily be used because it relies on tools and concepts familiar to as-

tronomers. A very few attempts have been made (see references in Roberts & Haynes 1994), but, to our knowledge, have not yielded a new classification. It has the power to determine the most discriminant characters for an objective classification.

Cladistics for astrophysics (astrocladistics) is probably the most promising approach. Basically, the only requirement for cladistics to be applicable and useful is the presence of a hierarchical organization of the diversity due to evolution (see a discussion in Brower 2000). Because galaxies are made of many different evolving processes, because galaxies evolve strongly through interactions, diversity is increasing with time and very probably organizes itself in a hierarchical way. (Note that “hierarchy” is here taken in the whole complexity of a galaxy, not only in mass or size.) Astrocladistics is the tool that will help us finding the hierarchical classification of galaxies.

Astronomers are particularly well equipped for cladistics. In principle, for each character, two states have to be defined: one will be said “ancestral” and the other one “derived”. In practice, it is possible to define more than two states (8 with current softwares), hence hypothesizing an evolutionary scheme for each character. Processes occurring in galaxies are relatively well known and can be modeled sometimes with great details. Observations are providing more and more detailed data, at larger and larger redshifts. Hence, character coding in astrophysics does not seem to be too much of a problem.

3. Application to the dwarf galaxies of the Local Group

We are in the process of applying astrocladistics to a sample of dwarf galaxies of the Local Group. Data are taken from Mateo (1998). There are 36 galaxies and 28 characters (properties). Each character is binned in up to 8 states, and these states are related to each other by a linear transformation series. This hypothesis is obviously debatable and will have to be made more sophisticated in subsequent studies. Among the 28 characters, some will certainly appear to be irrelevant after a complete analysis is performed. This work, still in progress, will be published elsewhere. But to illustrate the power of astrocladistics, we present what kind of results and interpretations can be obtained.

At the end of the analysis, groups (or clades) of galaxies are found on the cladogram (Fig. 1). This provides the classification. Galaxies in each group share a common history: it can be the same kind of past encounters, of environment during their lives, etc... In particular, one of the group might well be tidal dwarfs. To understand the history, it is necessary to look at which character changes on which branch of the

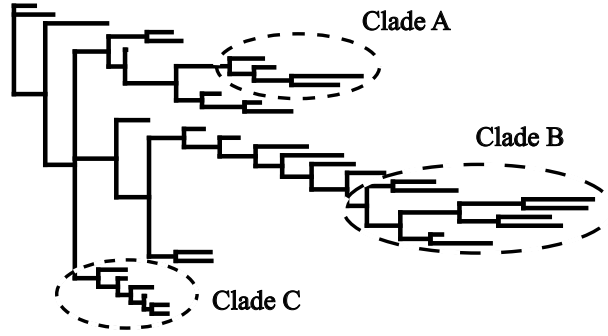


Figure 1.

cladogram. For instance, galaxies of the three clades defined on Fig. 1 have the following properties:

Clade A	Clade B	Clade C
spheroidal	irregular	spheroidal
low ellipticity	high ellipticity	high ellipticity
low HI mass	high HI mass	low HI mass
low Fe/H	all Fe/H	high Fe/H
low M/L _*	low M/L _*	high M/L _*

It is clear from the table that it is inappropriate to name these galaxies after their morphology or one other property, like birds cannot be described only by wings and mammals only by warm blood. A new taxonomy is to be invented, but this is not a trivial problem.

References

- Brower, A.: 2000, 'Evolution is not a necessary assumption of cladistics', *Cladistics* **16**, pp. 143-154
 Robert, M.S., Haynes, M.P.: 1994, 'Physical parameters along the Hubble sequence', *ARAA* **32**, pp. 115-152
 Stewart, C.-B.: 1993, 'The powers and pitfalls of parsimony', *Nature* **361**, pp. 603-607
 Mateo, M.: 1998, 'Dwarf galaxies of the Local Group', *ARAA* **36**, pp. 435-506